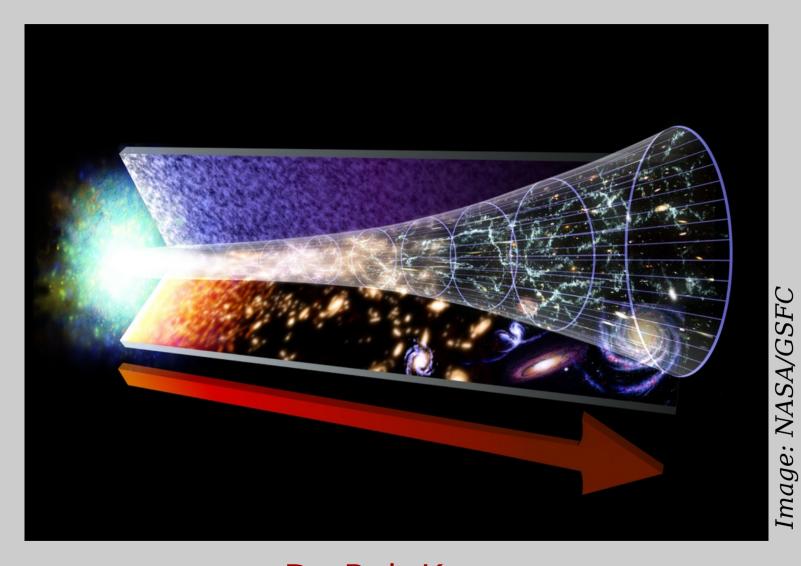
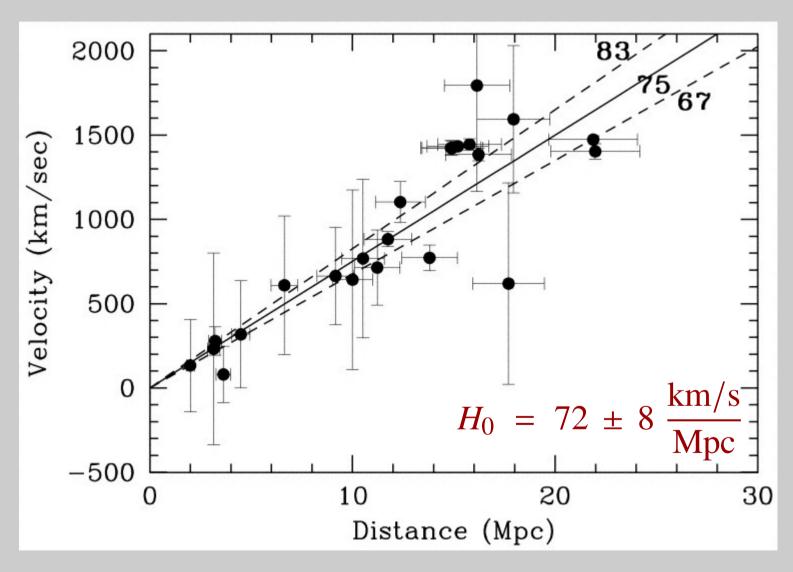
A big outstanding question in cosmology: Tension in the Hubble Constant



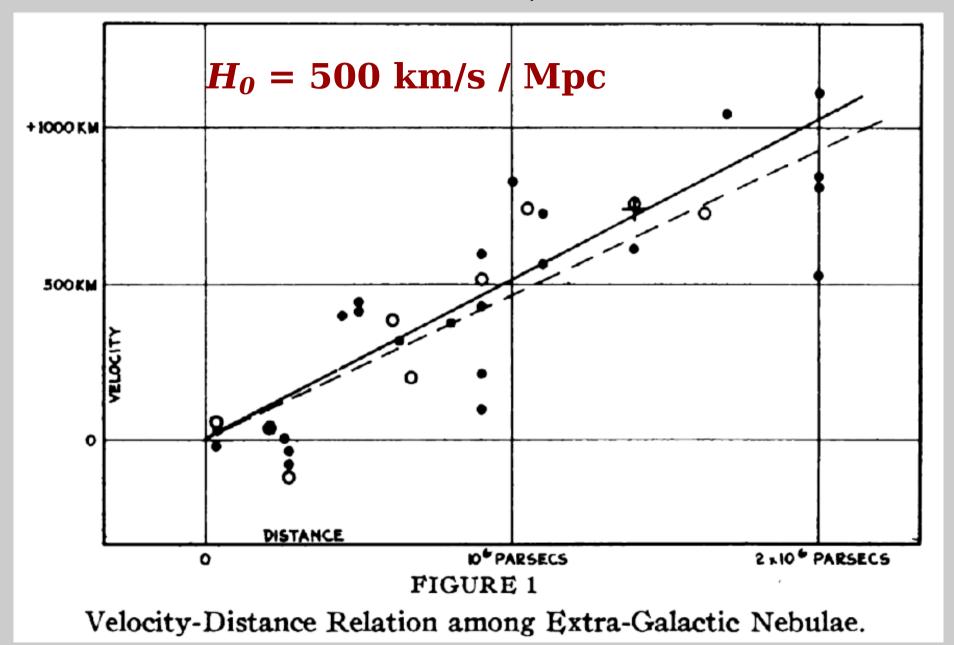
Dr. Rob Knop Science Circle, Second Life, 2019-09-28

The Hubble Constant is the current expansion rate of our Universe



Freedman et al., 2001, ApJ, 553, 47

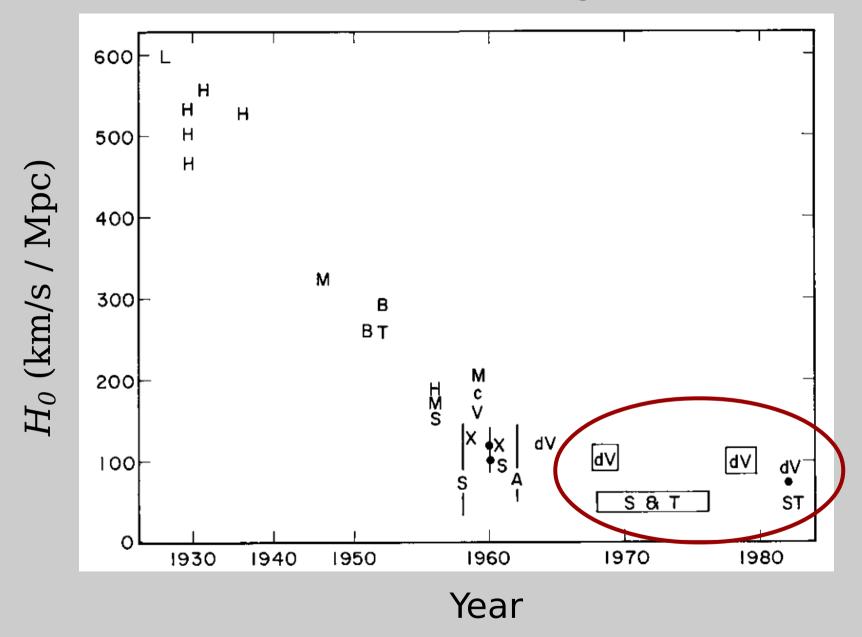
Edwin Hubble, 1929



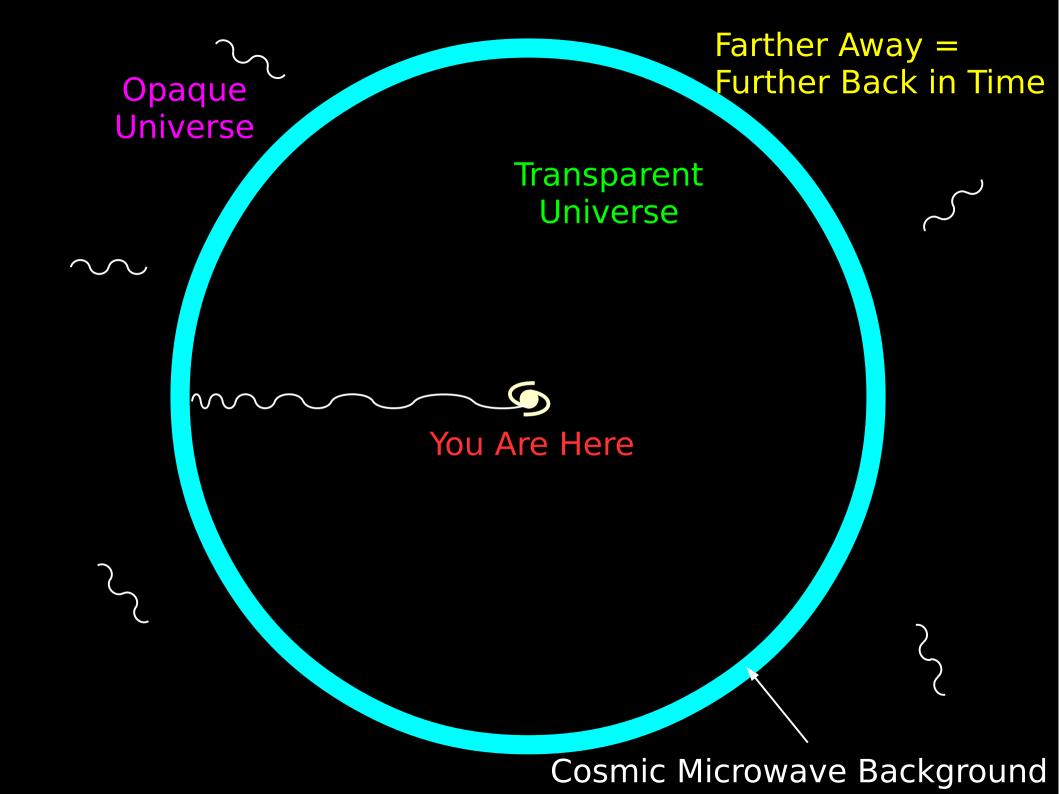
(Note: Lemaître 1927, Freedman 1922, et al.)

Hubble, 1929, PNAS, 15, 168

Hubble Constant Estimates through the mid-80s



Trimble, 1997, Space Science Rev. 79, 793 Available CC-BY-4.0 thanks to UC Irvine's eScholarship program



COBE: The Cosmic Background Explorer

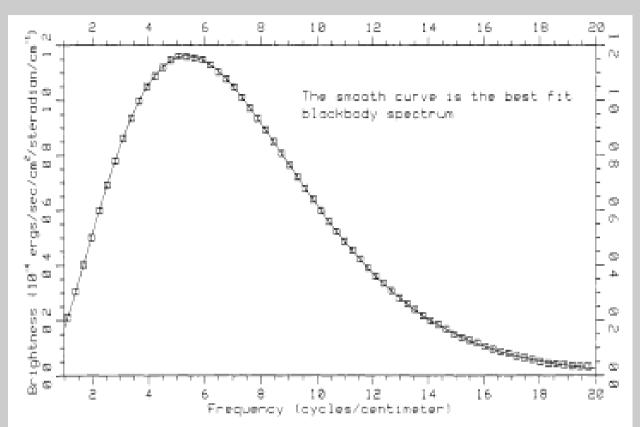
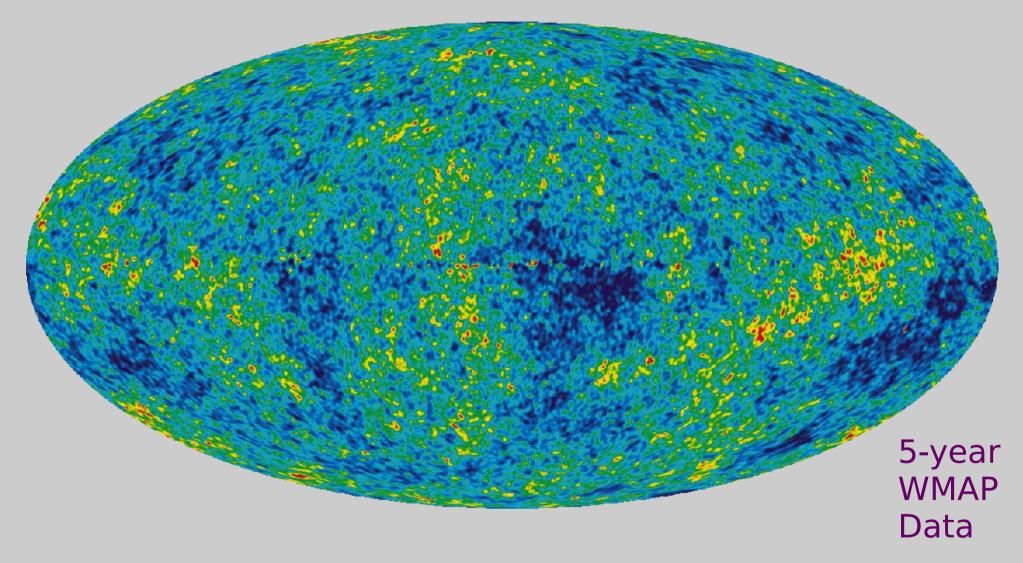


Fig. 2.—Preliminary spectrum of the cosmic microwave background from the FIRAS instrument at the north Galactic pole, compared to a blackbody. Boxes are measured points and show size of assumed 1% error band. The units for the vertical axis are 10^{-4} ergs s⁻¹ cm⁻² sr⁻¹ cm.

Mather et al., 1990, ApJ, 354, L37

- The spectrum of the CMB perfectly matches a blackbody at 2.74K
- This temperature is consistent across the sky to 1 part in 1000

CMB Anisotropy 2: Fluctuations



Maximum fluctuation amplitude : 75 μK

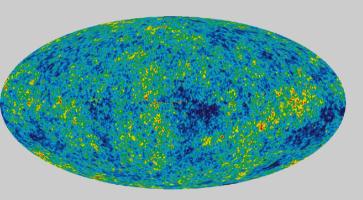
(The CMB is smooth to one part in 40,000)

Bennet et al., 2014, ApJ, 749, 135 "The 1% Concordance Hubble Constant"



Distance Ladder Measurements

$$H_0 = 73.0 \pm 2.4 \text{ km/s} / \text{Mpc}$$



CMB Measurements

WMAP: $H_0 = 68.76 \pm 0.84 \text{ km/s} / \text{Mpc}$

Planck: $H_0 = 67.3 \pm 1.2 \text{ km/s} / \text{Mpc}$

Concordance Value : $69.6 \pm 0.7 \text{ km/s} / \text{Mpc}$

< 20 tension

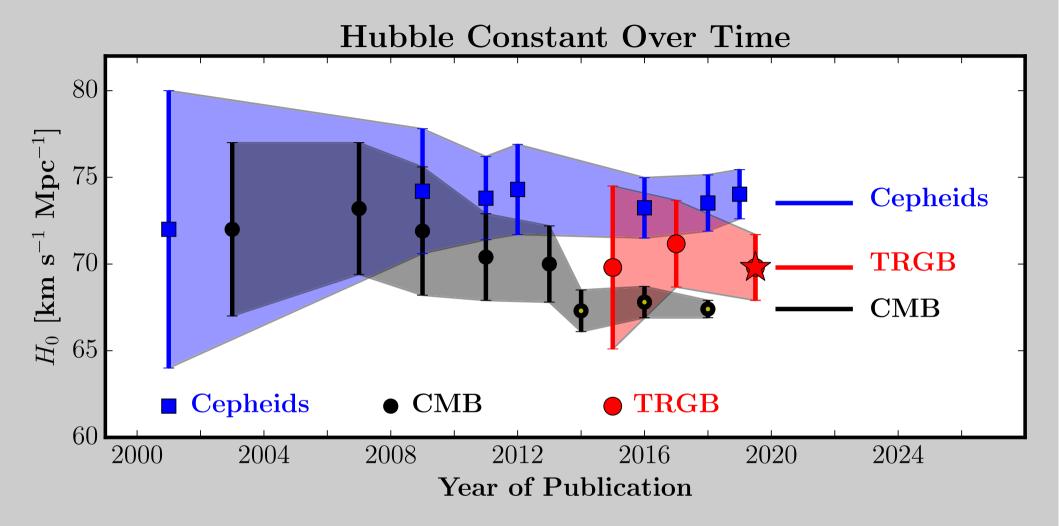
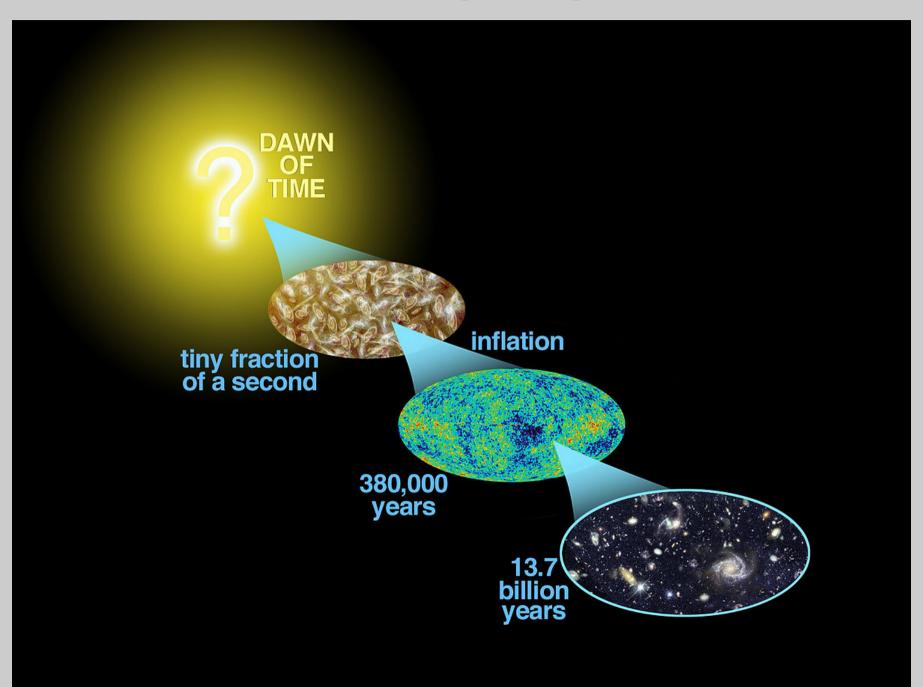


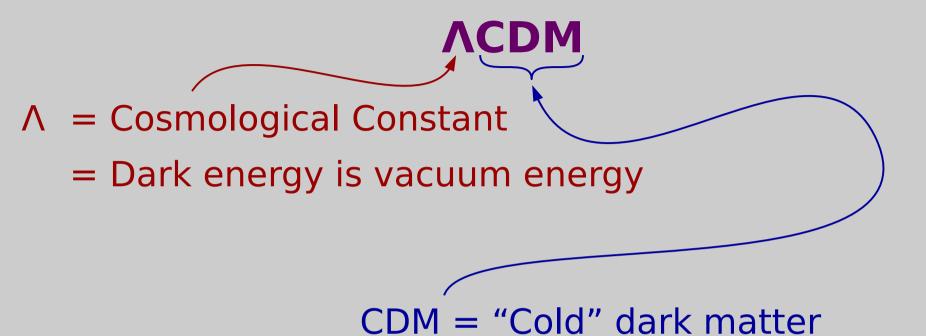
Image: NASA/WMAP Science Team

Our fundamental model of the history of our Universe:

The Big Bang



Our current specific Big Bang model:



= Dark matter particles are blobbed around galaxy clusters; particles move much less than the speed of light

Resolving the Hubble Tension

- Unidentified systematics in one or more methods for determining H0
- ACDM isn't quite right (1) dark energy is more complicated than vacuum energy
- ACDM isn't quite right (2) Dark Matter isn't strictly cold, or interacts in ways we don't know
- We live in a local low-density region that affects "nearby" measurements of H0
- Something else